# Light-Shifts of an Integrated Filter-Cell Rubidium Atomic Clock

May 25, 2015

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Prepared for:

Space and Missile Systems Center Air Force Space Command 483 N. Aviation Blvd. El Segundo, CA 90245-2808

Contract No. FA8802-14-C-0001

Authorized by: Engineering and Technology Group



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1. REPORT DATE 25 MAY 2015		2. REPORT TYPE <b>Final</b>		3. DATES COVERED -		
4. TITLE AND SUBTITLE  Light-Shifts of an Integrated Filter-Cell Rubidium Atomic			ic Clock	5a. CONTRACT NUMBER <b>FA8802-14-C-0001</b>		
				5b. GRANT NUMBER		
				5c. PROGRAM ELEMENT NUMBER		
6. AUTHOR(S)				5d. PROJECT NUMBER		
James C. Camparo	0		5e. TASK NUMBER			
			5f. WORK UNIT NUMBER			
	IZATION NAME(S) AND A rporation 2310 E. E	Segundo, CA	8. PERFORMING ORGANIZATION REPORT NUMBER TOR-2015-02236			
<b>Space and Missile</b>	ORING AGENCY NAME(S)  Systems Center Air  Second of CA 20245	10. SPONSOR/MONITOR'S ACRONYM(S)  SMC				
Aviation Bivd. El S	Segundo, CA 90245		11. SPONSOR/MONITOR'S REPORT NUMBER(S)			
12. DISTRIBUTION/AVAI Approved for pub	LABILITY STATEMENT lic release, distribut	ion unlimited				
13. SUPPLEMENTARY NO The original docum	OTES ment contains color	images.				
14. ABSTRACT						
15. SUBJECT TERMS						
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**Report Documentation Page** 

Form Approved OMB No. 0704-0188

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## Light-Shifts of an Integrated Filter-Cell Rubidium Atomic Clock

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#### Abstract

In this work, we consider the light-shift coefficients observed in rubidium (Rb) atomic frequency standards, and how the units of measurement reported for light-shift coefficients trace back to the underlying physical parameters of the light-shift. We then measure the light-shift coefficient in an integrated filter-cell, commercial Rb clock, which is similar to the Rb clock design currently flying on Galileo and BeiDou satellites. We measure the light-shift coefficient for two different rf-discharge lamps (i.e., a pure <sup>87</sup>Rb lamp and a lamp filled with the natural Rb isotope abundance), and show that under certain conditions it is possible to combine light-shift data. Our light-shift coefficients are consistent with those reported for the Galileo Rb clock under the assumption of a natural (or <sup>85</sup>Rb isotopically enriched) rf-discharge lamp for the Galileo clock.

#### I. Introduction

#### Overview

The rubidium (Rb) atomic frequency standard is the workhorse of precise atomic timekeeping in space due to its low weight, small volume, and low power consumption [1]. Moreover, the device has excellent frequency stability out to (and beyond) 10<sup>4</sup> seconds averaging time, rivaling the performance of passive hydrogen masers [2,3]. However, one of the clock's more significant issues is long-term frequency drift, which may be related to the light-shift effect [4,5]: a basic physics atomic perturbation that causes the Rb clock's output frequency to depend on the intensity of (atomic-signal-producing) lamplight. Moreover, on-orbit Rb atomic clocks have displayed frequency jumps that are correlated with lamplight jumps [6,7]. Thus, understanding the light-shift effect has both basic and applied physics implications.

Here we discuss our measurements of the light shift effect in a commercial, integrated filter-cell [8], Rb atomic clock, similar to the Rb atomic clock flying on Galileo GNSS satellites [9]. As illustrated in Fig. 1, the physics package of the prototypical Rb clock (i.e., the separated filter-cell clock) is composed of a lamp, a filter cell, a resonance cell, and a photodetector, with the resonance cell inside a microwave cavity that is tuned to the 87Rb atom's hyperfine resonance frequency of 6834.7 MHz. The lamplight passes through the filter cell, and in so doing has its optical spectrum "shaped" so that it can efficiently produce an atomic signal via the process known as optical pumping [10,11]. As a consequence of optical pumping, the steady-state population of atoms in the absorbing hyperfine level (i.e., F=1) is reduced, and so the Rb vapor in the resonance cell becomes (to some degree) transparent to the lamplight. With the Rb vapor absorbing little lamplight, more light passes through to the photodiode.

If microwaves are applied to the vapor at the hyperfine resonance frequency, atoms are forced by the microwave signal to return to the absorbing hyperfine state. Consequently, more lamplight gets absorbed by the vapor, and less passes through to the photodiode. The change in transmitted light corresponds to the amplitude of the atomic clock signal. The magnitude of this change will, of course, depend on the efficiency of the optical pumping process, which ultimately traces back to the filter cell's ability to optimally shape the lamplight's spectrum.

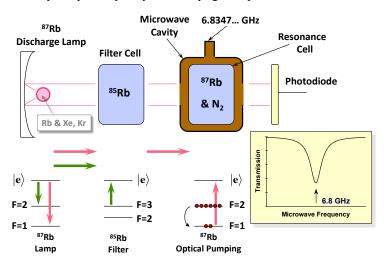


Figure 1: Block diagram of the prototypical separated filter-cell rubidium (Rb) atomic clock. The <sup>87</sup>Rb lamp emits spectral lines that originate in a rubidium excited state and terminate on one of the two ground-state hyperfine levels of the atom. Due to a coincidence of nature, one of these spectral lines is absorbed by the <sup>85</sup>Rb atoms in the resonance cell, allowing the other spectral line to pass through, where it can create a population imbalance between the ground-state hyperfine levels via optical pumping.

In the 1970s, Efratom, a Rb clock manufacturing company, commercialized the integrated filter-cell Rb clock [12]. For an integrated filter-cell clock, the resonance cell and filter cell are combined in one: the Rb vapor in the filter/resonance cell combination contains both rubidium isotopes. The filtering action of the lamplight by <sup>85</sup>Rb atoms occurs in the front of the filter/resonance cell, so that the <sup>87</sup>Rb atomic signal is generated towards the back of the filter/resonance cell. An advantage to the integrated filter-cell approach is that one set of heaters is eliminated from the clock's design (*i.e.*, the separate filter-cell heater), which reduces the clock's overall power budget. We note somewhat parenthetically that the first Rb clocks flown on GPS were of an Efratom integrated filter-cell design [13], and integrated filter-cell Rb clocks have been flown on Galileo [9] and BeiDou satellites [14].

## II. The Light-Shift

General Considerations

In a semiclassical formalism, the light-shift arises as a second-order interaction between an atom's induced dipole moment and the perturbing electrical field [15,16]. (For a classical description of the light shift, see the Appendix.) If we imagine the atom as a polarizable medium as illustrated in Fig. 2, with a (frequency dependent) polarizability  $\alpha(\omega)$ , then the perturbing electric field of the light,  $\bar{E}(\omega)$ , will induce a dipole moment  $\bar{p}(\omega)$  in the atom:

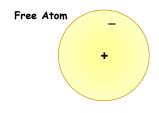
$$\vec{p}(\omega) = \alpha(\omega)\vec{E}(\omega). \tag{1}$$

This dipole will interact with the electric field that produced it, so that to *second-order* in the electric field strength there is an interaction energy between the atom and the light,  $\Delta \epsilon$ :

$$\Delta \varepsilon = -\frac{1}{2} \vec{p}(\omega) \cdot \vec{E}(\omega). \tag{2}$$

This interaction perturbs the atom's energy level structure, giving rise to a shift in the atom's ground-state hyperfine splitting,  $h\nu_{hfs}$ :

$$\frac{\delta f}{f_o} = \frac{\Delta \varepsilon}{h v_{hfs}} = -\frac{\vec{p}(\omega) \cdot \vec{E}(\omega)}{2h v_{hfs}} = -\left(\frac{1}{2h v_{hfs}}\right) \alpha(\omega) |\vec{E}(\omega)|^2 \cdot (3)$$



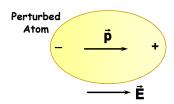


Figure 2: Illustration of an atom as a polarizable medium. In the presence of an electric field a dipole moment is induced in the atom, which can then interact with the electric field to produce a perturbation of the atom's energy level structure.

#### The Light-Shift Coefficient

The light-shift as expressed by Eq. (3) is valid for a monochromatic light wave. However, as is well known the lamplight's spectrum after filtering can be fairly broadband and complex [17,18], and each spectral component of the lamplight will contribute to the overall light-shift,  $\Delta y_{LS}$ , via Eq. (3). Consequently, if we represent the normalized intensity spectrum of the lamplight (after filtering) in the frequency range  $\omega$  to  $\omega$ +d $\omega$ , by S( $\omega$ )d $\omega$ :

$$\int_{0}^{\infty} S(\omega) d\omega = 1, \tag{4}$$

and take  $I_o$  as the total Rb light reaching the  $^{87}\text{Rb}$  atoms in the atomic-signal generating region of the clock's physics package, then

$$\Delta y_{LS} = I_o \int_0^\infty \alpha(\omega) S(\omega) d\omega_o = \beta_{LS} I_o, \qquad (5)$$

where  $\beta_{LS}$  is the light-shift coefficient of the clock. Clearly,  $\beta_{LS}$  will depend on the emission spectrum of the lamp as well as the filtering action in the clock.

Equation (5) is valid for the clock's full light-shift. However, we are typically not interested in the full light-shift, but rather variations in the light-shift about its average value:  $\delta[\Delta y_{LS}],$  which for convenience we will write simply as  $\delta_{LS}.$  Therefore, defining  $\langle I_o \rangle$  as the average light intensity emitted by the lamp we have

$$\delta_{LS} \equiv \delta[\Delta y_{LS}] = \beta_{LS} (I_o - \langle I_o \rangle) = \beta_{LS} \langle I_o \rangle \frac{\Delta I}{\langle I_o \rangle} = \kappa_{LS} \frac{\Delta I}{\langle I_o \rangle}. (6)$$

 $\kappa_{LS}$  is the parameter routinely reported in the literature for a clock's light-shift coefficient, typically in units of %-1. However, as can readily be seen from Eq. (6), clock-to-clock comparisons among  $\kappa_{LS}$  values should be treated with caution [19], since  $\kappa_{LS}$  depends on  $\langle I_o \rangle$ . The parameter of real comparative value is  $\beta_{LS}$ .

## Comparing $\kappa_{LS}$ Values

While caution must be exercise in comparing  $\kappa_{LS}$  values among clocks and clock operating conditions, there are certain situations where comparison may be possible. Specifically, let's assume that the following conditions hold for two Rb atomic clocks: Clock #1 and Clock #2

- a. The filtering action of the clocks is similar. Specifically, the two clocks are both either separated filter cell or integrated filter cell. Additionally, we assume that the optical depth of the <sup>85</sup>Rb filtering vapor is similar in the two clocks.
- b. The rf-discharge lamps of the two clocks operate with similar electronics, so that the degrees of plasma ionization and the electron temperatures are not too dissimilar [20].
- c. The lamps of the two clocks have the same Rb isotope composition (e.g., pure  $^{87}$ Rb or natural Rb 72.2%  $^{85}$ Rb and 27.8%  $^{87}$ Rb).
- d. The lamps operate at not too dissimilar temperatures.

Given these assumptions,  $\beta_{LS}$  for the two clocks should be the same, leaving only  $\langle I_o \rangle$  to differ. Considering Eq. (6) for the two clocks, we then have

$$\delta_{LS}(l) = \kappa_{LS}(l) \frac{\Delta I}{\langle I_o \rangle} = \beta_{LS} \langle I_o \rangle_2 \left( \frac{\langle I_o \rangle_1}{\langle I_o \rangle_2} \right) \frac{\Delta I}{\langle I_o \rangle}, \quad (7a)$$

$$\frac{\delta_{LS}(1)}{\delta_{LS}(2)} = \frac{\kappa_{LS}(1)}{\kappa_{LS}(2)} = \left(\frac{\langle I_o \rangle_1}{\langle I_o \rangle_2}\right),\tag{7b}$$

which provides a means of comparing  $\kappa_{LS}$  values. (Here, we have assumed that any differences in the lamps' light intensities affects the numerator and denominator of  $\Delta I/\langle I_o \rangle$  equivalently.)

We can go a bit further with this line of reasoning. Specifically,  $\langle I_o \rangle$  should primarily depend on the vapor density of rubidium in the lamp [21]. Since Rb vapor density depends exponentially on the temperature of the liquid Rb pool in the lamp, all other things being equal the ratio of  $\langle I_o \rangle$  for lamps #1 and #2 should be given by

$$\frac{\left\langle I_{o}\right\rangle_{1}}{\left\langle I_{o}\right\rangle_{2}} = \operatorname{Exp}\left[\frac{\Delta G}{k_{B}}\left(\frac{1}{T_{2}} - \frac{1}{T_{1}}\right)\right],\tag{8}$$

where  $\Delta G$  is the energy of Rb vaporization, which has the value  $1.313\times10^{-12}$  ergs [22],  $k_B$  is Boltzmann's constant, and  $T_J$  is the base temperature of lamp #J. Using Eq. (8) in Eq. (7b), we can then write

$$\frac{\delta_{LS}(1)}{\delta_{LS}(2)} = \frac{\kappa_{LS}(1)}{\kappa_{LS}(2)} = \exp\left[\frac{\Delta G}{k_B}\left(\frac{1}{T_2} - \frac{1}{T_1}\right)\right]$$
(9)

## III. Experiment

There are two ideas worth testing from the above: 1)  $\kappa_{LS}$  is strongly affected by the rf-discharge lamp's emission spectrum, and 2)  $\delta_{LS}$  values are comparable (under appropriate conditions) by the simple scaling of Eq. (7b).

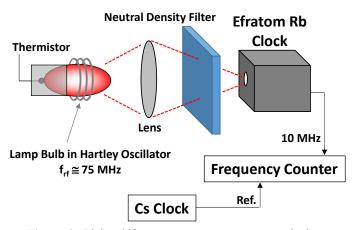
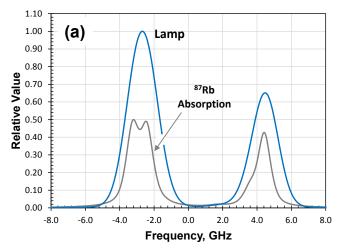


Figure 3: Light shift measurement apparatus employing a commercial, integrated filter-cell Rb atomic clock.

For our experiment, we employed a commercial FRK Rb atomic clock manufactured by Efratom [12], which has the advantage that the lamp bulb can be easily removed from the unit allowing a clear optical path from the outside to the clock's inner physics package. Using a separate external lamp for optical pumping and clock signal detection, as illustrated in Fig. 3, we placed neutral density filters in front of the lamp and recorded the Rb clock's frequency relative to a cesium (Cs) atomic clock.



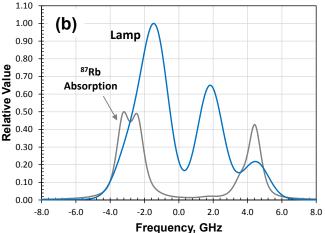


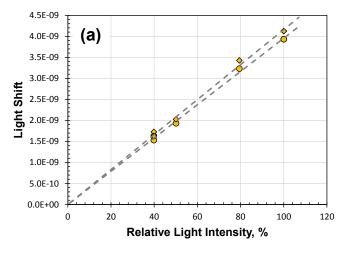
Figure 4: (a) Nominal spectrum form a pure <sup>87</sup>Rb lamp. (b) Nominal spectrum from a natural Rb lamp. For these calculated spectra we assumed a 1.6 GHz wide lamp line [23], and 40 torr of Xe as a buffer gas in the clock's resonance cell. (The actual choice of buffer gas for this figure is not particularly important.) For the pure <sup>87</sup>Rb lamp, most of the spectral intensity (*i.e.*, the lamp's "spectral weight") falls at frequencies *outside of* the <sup>87</sup>Rb hyperfine splitting; for the natural Rb lamp, most of the spectral intensity falls at frequencies *within* the hyperfine splitting.

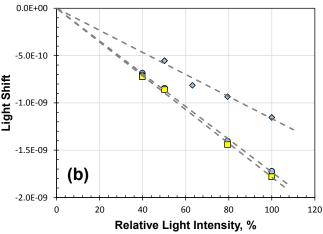
We employed three lamps in our study, two contained pure <sup>87</sup>Rb while the third contained natural Rb. Figures 4a and 4b illustrate the different emission spectra expected for these lamps. The lamps were operated at 137 °C and 125 °C. For the lower temperature, optical elements were added to the simplified schematic shown in Fig. 3 in order to collect enough light to operate the clock. As a consequence, our results are not expected to follow Eq. (9) for the two lamp operating temperatures. However, if intensity scaling is viable for similar lamps operating under different temperatures, then our results should follow the more general Eq. (7b).

## IV. Results

Figures 5a and 5b show the raw light-shift results: Fig. 5a corresponds to the natural Rb lamp, while Fig 5b shows the results for our two pure <sup>87</sup>Rb lamps. Taken in combination with

Figs. 4, these results demonstrate the dependence of  $\kappa_{LS}$  on lamp spectrum. Specifically, with the natural lamp's "spectral weight"\* within the <sup>87</sup>Rb hyperfine splitting we expect (and find)  $\kappa_{LS} > 0$ ; with the pure <sup>87</sup>Rb lamp's spectral weight outside of the <sup>87</sup>Rb hyperfine splitting we expect (and find)  $\kappa_{LS} < 0$  [1].

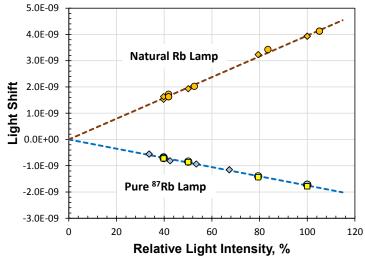




*Figure 5*: (a) Raw light-shift data for the natural Rb lamp: circles correspond to  $T_{lamp} = 137$  °C and diamonds correspond to  $T_{lamp} = 125$  °C. (b) Raw light-shift measurements for our two pure  $^{87}$ Rb lamps: circles correspond to Lamp #1 at 137 °C; diamonds correspond to Lamp #1 at 125 °C, and squares correspond to Lamp #2 at 137 °C.

In Fig. 6, we have rescaled the raw light-shift results according to Eq. (7b). Specifically, for the natural Rb lamp we used linear regression of the raw data to compute an estimated light-shift at full light intensity for each of the temperatures. We call these estimated light-shifts  $\hat{\delta}_{LS}(137)$  and  $\hat{\delta}_{LS}(125)$ . We then multiplied the raw light intensities for the  $T_{lamp} = 125$  °C data by the ratio  $\hat{\delta}_{LS}(125)/\hat{\delta}_{LS}(137)$ . In effect, rather than changing the

light-shift values, we simply repositioned them along the light intensity axis. Saying this differently, we identified the 125 °C light-shifts with the normalized light intensities that they would have corresponded to for the 137 °C experiment. A similar procedure was employed for the pure <sup>87</sup>Rb lamp, except that all light intensities were repositioned based on Lamp #1's 137 °C values.



*Figure 6*: Rescaled light-shift measurements. Symbols are as in Fig. 5.

As Fig. 6 clearly shows, all the light-shift results for a *single type* of lamp fall on one line, which demonstrates that it is possible to rescale  $\kappa_{LS}$  values for similar lamps. Taken together, we find for the natural Rb lamp  $\kappa_{LS}=+4.0\times10^{-11}/\%$ , and for the pure  $^{87}\text{Rb}$  lamp  $\kappa_{LS}=-1.8\times10^{-11}/\%$ . As noted previously, the signs of these  $\kappa_{LS}$  values are consistent with the lamps' spectral weights. Moreover, the relative magnitude of these  $\kappa_{LS}$  values is consistent with the fact that relatively more spectral intensity falls *within* the hyperfine splitting for the natural lamp than falls *outside of* the hyperfine spitting for the pure  $^{87}\text{Rb}$  lamp.

## V. Summary and Discussion

In this work, we have considered the light-shift coefficient of an integrated filter-cell Rb atomic clock operated with two different Rb rf-discharge lamps: a pure <sup>87</sup>Rb lamp and a natural Rb lamp (*i.e.*, 72.2% <sup>85</sup>Rb and 27.8% <sup>87</sup>Rb). As anticipated, the sign of the light-shift coefficient followed from a simple inspection of the spectral weight of the lamp's emission spectrum. Though the filtering action of the lamplight has the potential to significantly affect the light-shift's sign (since it is the *filtered* lamplight that ultimately determines the light-shift), in the present integrated filter-cell clock experiments the filtering was clearly not so strong as to significantly modify the effect of the lamp's spectral weight.

<sup>\*</sup> As discussed in the caption of Fig. 4, we define the spectral weight of a lamp as the excess of intensity falling either outside the <sup>87</sup>Rb hyperfine splitting or within that splitting. Specifically, for the pure

<sup>&</sup>lt;sup>87</sup>Rb lamp 6% more light falls outside of the hyperfine splitting than within; for the natural Rb lamp, 67% more light falls within the hyperfine splitting than outside of that splitting.

It is interesting to note that the magnitudes of the light-shift coefficients that we obtained (even in units of %-1) are not outside the range of expectation:  $\sim 10^{-11}$ /%. Specifically, it has been noted that the light-shift coefficient for the separated filter-cell, Block-IIR GPS Rb clock is  $-1.4 \times 10^{-12}$ /% [7], while Droz et al. [24] have reported a light-shift coefficient for Galileo's integrated filter-cell clock of +1.5×10<sup>-10</sup>/%. Additionally, Bloch et al. [25] provide evidence for a light-shift coefficient of -2.2×10<sup>-10</sup>/% for the Milstar Rb clock, which is a separated filter-cell design [26]. If we assume that moving the lamp outside of the atomic clock reduces the intensity of the lamplight actually reaching the atoms in the filter/resonance cell by a factor of roughly 10, which is certainly reasonable, than our integrated filter-cell light-shift coefficients are perfectly in line with those of Droz et al.... assuming, of course, that the Galileo Rb atomic clock operates with a natural (or 85Rb isotopically enriched) rf-discharge lamp.

## Acknowledgement

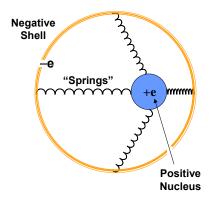
The author would like to thank Dr. Robert Frueholz for his help in performing these experiments.

## Appendix: Classical Description of the Light Shift

Though the light-shift is a consequence of an atom's quantum mechanical nature, the mechanism of the light-shift can nonetheless by understood classically [15]. As illustrated in Fig. A.1, we imagine the "classical atom" as a positive nucleus attached to a negatively charged shell with springs of force constant k [27]. Defining x(t) as the deviation of a spring from its equilibrium position, the equation of motion for this atomic oscillator is just that of a damped, driven harmonic oscillator:

$$\frac{d^2x(t)}{dt^2} + 2\gamma \frac{dx(t)}{dt} + \frac{k}{m}x(t) = -\frac{e}{m}E(t). \tag{A.1}$$

Here,  $2\gamma$  is a damping coefficient, kx is the restoring force on the spring, m is the mass of the negative shell (*i.e.*, the electron mass) and E(t) is the instantaneous value of an electric field impinging on the classical atom.



<sup>†</sup> The light-shift coefficient was determined by examining the correlated change in clock frequency and lamplight when the vacuum system for the Rb clocks failed: See Figs. 2a and 2b in Ref. 25.

Figure A.1: The "classical atom," where the atom's negatively charged shell is attached to a positively charged nucleus via mechanical springs.

To proceed, we consider a monochromatic field such that  $E(t) = E_0 \exp[i\omega t]$ , where  $\omega$  is the laser frequency, which allows us to write  $x(t) = x_0 \exp[i\omega t]$ . Using these expressions in Eq. (A.1), and defining the atom's resonant frequency,  $\omega_0$ , as  $\sqrt{k/m}$  we obtain

$$x_{o} = -\frac{\left(e/m\right)E_{o}}{\left(\omega_{o}^{2} - \omega^{2}\right) + i2\gamma\omega}.$$
 (A.2)

Since we are only interested in optical field frequencies very near  $\omega_0$ , we can expand  $\omega^2$  in a Taylor series about  $\omega_0$ :

$$\omega^2 = \omega_o^2 + (\omega - \omega_o) 2\omega_o + (\omega - \omega_o)^2, \qquad (A.3)$$

and only retain the first two terms on the right-hand side of Eq. (A.3). Further, under this resonance approximation we can replace  $\omega$  in the imaginary term appearing in the denominator of Eq. (2) with  $\omega_0$ , so that we finally obtain

$$x_o = -\left(\frac{e}{2m\omega_o}\right) \frac{E_o}{\left((\omega_o - \omega) + i\gamma\right)}.$$
 (A.4)

We now note that as the classical atom's shell oscillates relative to the location of the nucleus, an oscillating electric dipole moment will be created in the system, p(t):

$$p(t) = -ex(t) = -ex_o e^{i\omega t} = p_o e^{i\omega t}.$$
 (A.5)

Thus, using Eq. (A.4) we obtain

$$p_o = \left(\frac{e^2}{2m\omega_o}\right) \frac{E_o}{\left((\omega_o - \omega) + i\gamma\right)}.$$
 (A.6)

Of course, from classical electrodynamics we known that the induced dipole moment in a dielectric medium is just given by  $p(\omega) = \alpha(\omega)E(\omega)$ , where  $\alpha(\omega)$  is the medium's (frequency-dependent) polarizability. Using Eq. (A.6), we therefore obtain a classical expression for the polarizability of the atom:

$$\alpha(\omega) = \frac{\left(e^2 / 2m\omega_o\right)}{\left((\omega_o - \omega) + i\gamma\right)}.$$
 (A.7)

In order to make contact with the quantum mechanical properties of the atom, Eq. (A6) must be modified slightly in order to account for the fact that in a real atom a single energy eigenstate of the atom may be coupled with a large number of other eigenstates. We accomplish this by introducing the concept of the *oscillator strength*, f, so that Eq. (A.7) becomes [28].

$$\alpha(\omega) = \left(\frac{e^2 f}{2m\omega_o}\right) \left(\frac{(\omega_o - \omega) - i\gamma}{(\omega - \omega_o)^2 + \gamma^2}\right). \tag{A.8}$$

As is clear from Eq. (A.8)  $\alpha(\omega)$  is a complex quantity with both a real and imaginary part.

We now note that there is an interaction energy, V, between a dipole moment and an external electric field:

$$V = -\vec{p} \cdot \vec{E} = -\frac{1}{2}\alpha(\omega)|E_{o}|^{2}.^{\ddagger}$$
 (A.9)

Given our expression for  $\alpha(\omega)$ , the interaction energy is also found to be complex:

#### References

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$$Re[V] = \Delta \varepsilon \sim \left(\frac{e^2 f}{4m\omega_o}\right) \left(\frac{(\omega - \omega_o)I_o}{(\omega - \omega_o)^2 + \gamma^2}\right). \quad (A.10a)$$

$$Im[V] \sim \left(\frac{e^2 f}{4m\omega_o}\right) \left(\frac{\gamma \, I_o}{\left(\omega - \omega_o\right)^2 + \gamma^2}\right). \tag{A.10b}$$

The real part of the perturbation gives rise to the shift in the atom's energy levels (*i.e.*, the light shift), while the imaginary part acts like a decay term, causing transitions between energy levels. In these expressions,  $I_o$  is the light intensity, and it is worth noting that for a monochromatic field (as considered here) if  $\omega = \omega_o$  the light shift is zero.

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<sup>\*</sup> The factor of ½ comes from averaging the interaction energy over the period of the oscillating electric field.

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# Light-Shifts of an Integrated Filter-Cell Rubidium Atomic Clock

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REPORT TITLE

Light-Shifts of an Integrated Filter-Cell Rubidium Atomic Clock

TOR-2015-02236 PUBLICATION DATE SECURITY CLASSIFICATION UNCLASSIFIED

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